

# How is the ionosphere formed?

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a. Photoionization

*Ionization Rate =*

(radiation intensity) x (ionization cross section) x (density)

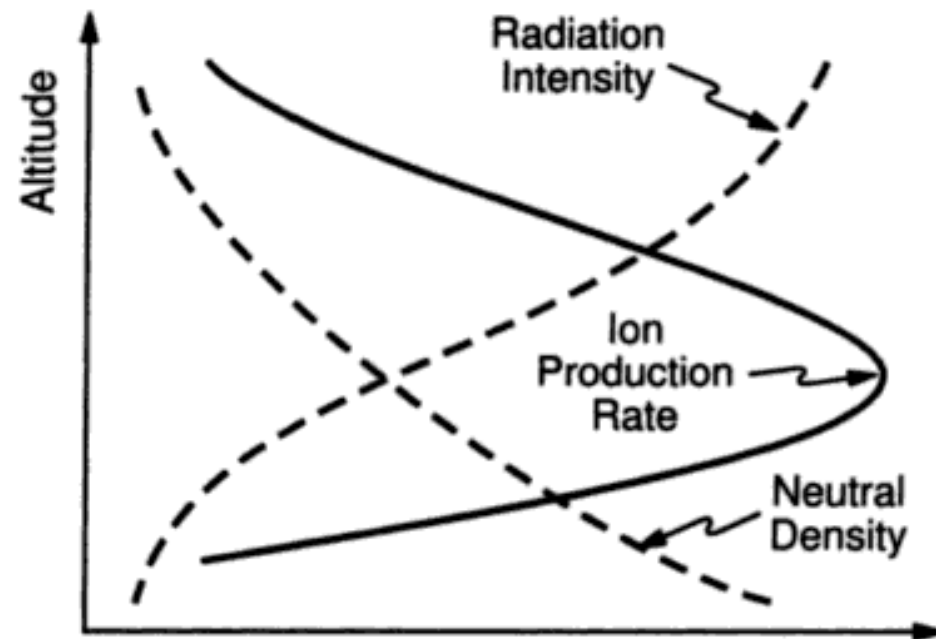
$$q_z = I_z \sigma_i n_z$$

$$q_z = I_\infty \sigma_i n_o \exp\left[-\frac{z - z_o}{H} - \frac{\sigma n_o H}{\mu} \exp\left[-\frac{z - z_o}{H}\right]\right]$$

Note:  $\sigma_i$  is ionization cross section and  $\sigma$  is absorption cross section. There is much wavelength and species dependent chemistry detail buried in  $\sigma_i$

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Formation of an ionized layer:



*Baumjohann & Treumann*

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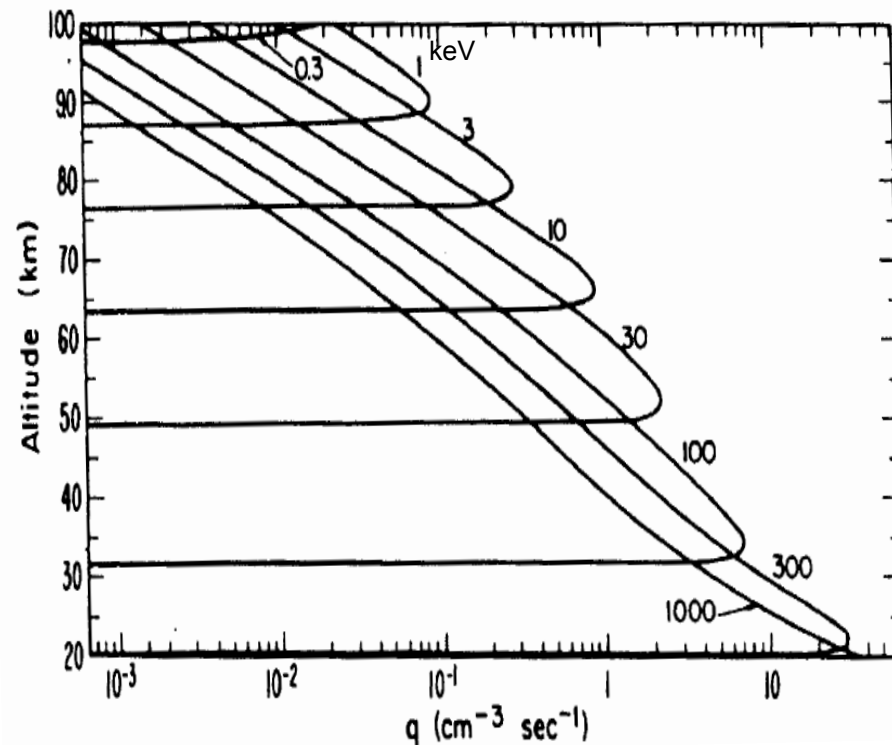
# How is the ionosphere formed?

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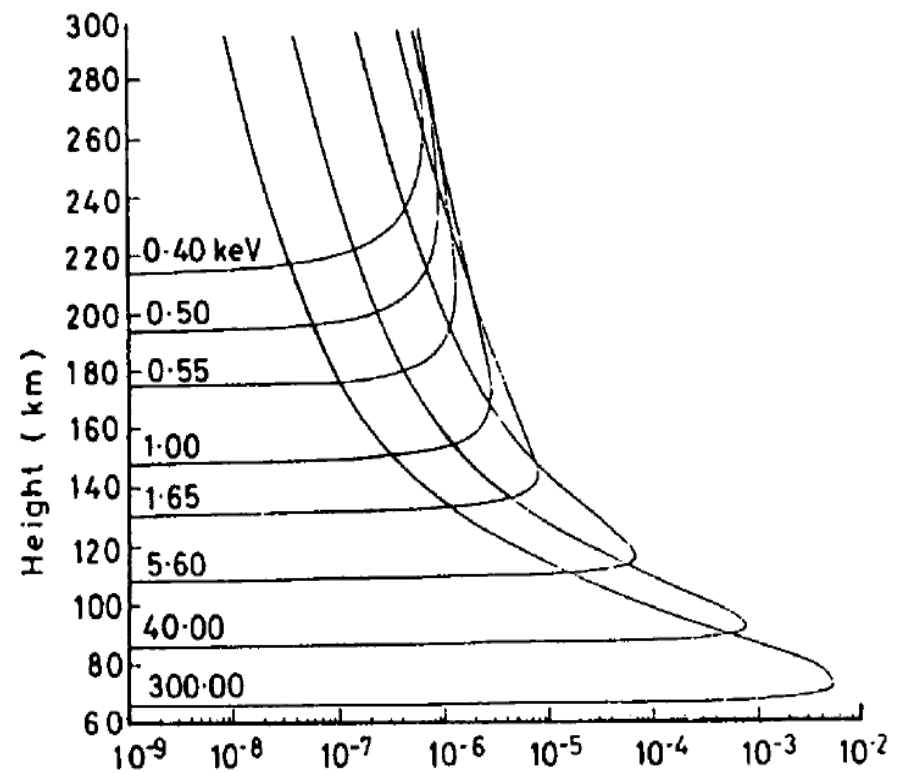
b. Impact ionization

Ionization Rate - Energy Dependant:

Due to proton precipitation



Due to electron precipitation



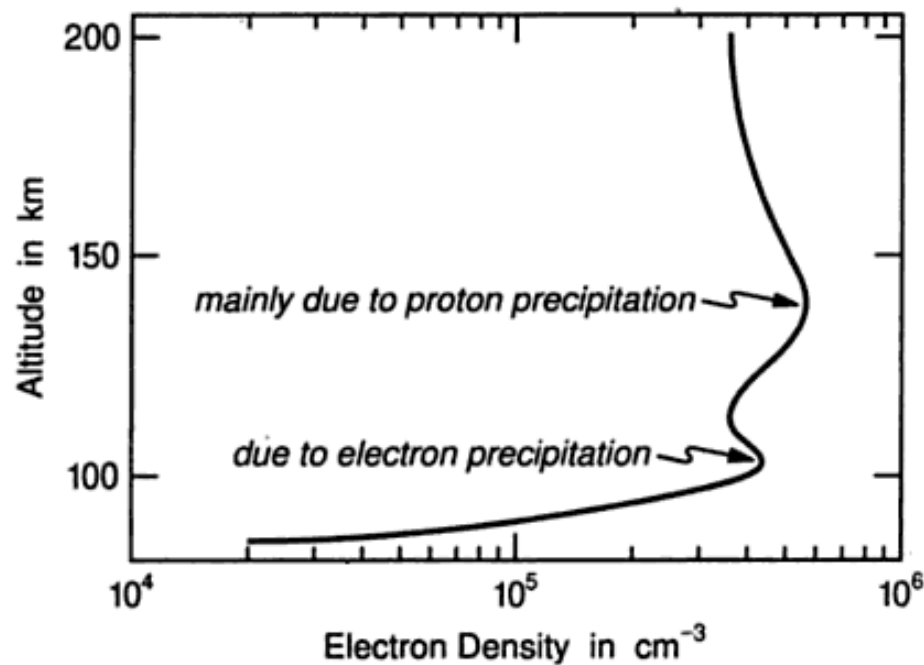
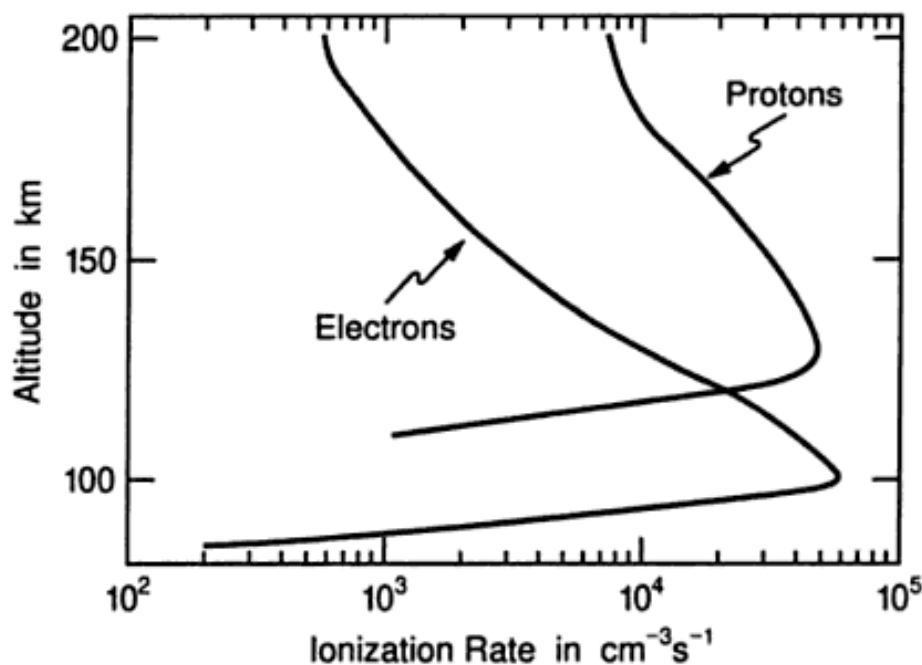
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b. Impact ionization

Ionization Rate Integrated over all precipitation energies:

Total ionization due to precipitation over diffuse aurora:



*Baumjohann & Treumann*

Note: These processes are most important at high latitudes (due to magnetic morphology), and at night when photoionization does not contribute!

# Structure of the Ionosphere

## Summing up:

Production of electrons from ionization of the neutral atmosphere from solar radiation and impacts from precipitating electrons and ions is counteracted by the processes of recombination and attachment.

$$\frac{dn_e}{dt} = q_{v,e} - \alpha_r n_e^2 - \beta_r n_e$$

Considering a system in equilibrium, i.e.  $dn_e/dt = 0$   
We find two simple proportionalities that determine the electron number density

# Structure of the Ionosphere

## Lower Ionosphere

In the lower ionosphere, recombination is the dominant loss process for electrons. Hence:

$$\frac{dn_e}{dt} = 0 \quad , \quad q_{v,e} - \alpha_r n_e^2 - \cancel{\beta_r n_e}^0 = 0$$
$$n_e = \left( \frac{q_{v,e}}{\alpha_r} \right)^{1/2}$$

At lower altitudes the electron density is proportional to the square root of the ionization rate.

# Structure of the Ionosphere

## Upper Ionosphere

In the upper ionosphere, attachment is the dominant loss process for electrons. Hence:

$$\frac{dn_e}{dt} = 0 \quad , \quad q_{v,e} - \cancel{\alpha n_e^2}^0 - \beta_r n_e = 0$$

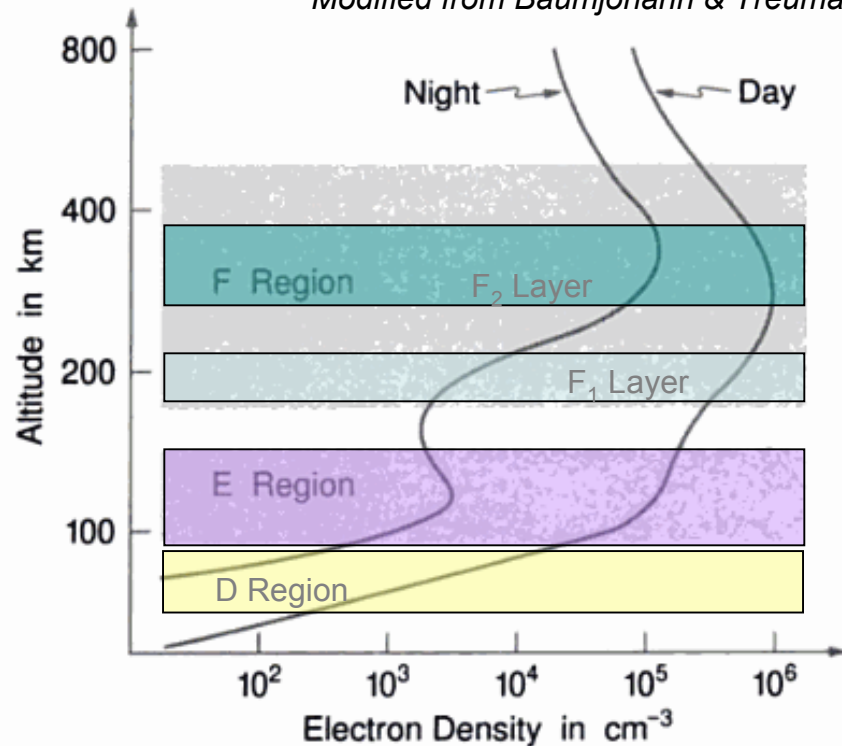
$$n_e = \frac{q_{v,e}}{\beta_r}$$

At higher altitudes the electron density is proportional to the ionization rate.

# Structure of the Ionosphere

## Layers of the Ionosphere:

*Modified from Baumjohann & Treumann*



**D-region:** Below 90 km, very weakly ionized and dominated by neutral gas dynamics (not a plasma)

**E-region:** Peaks at 110 km. Mainly produced by long wavelength UV (90nm) ionizing molecular Oxygen. At high latitudes impact ionization also plays important role. Electron density greatly reduced at night.

F-region: **F<sub>1</sub>** peaks at ~200 km, and is formed by shorter  $\lambda$  UV due to composition. Like the E-region, it has a strong diurnal signature.

**F<sub>2</sub>** peaks at ~300 km, at high altitudes only attachment counteracts ionization (which is limited by neutral density).