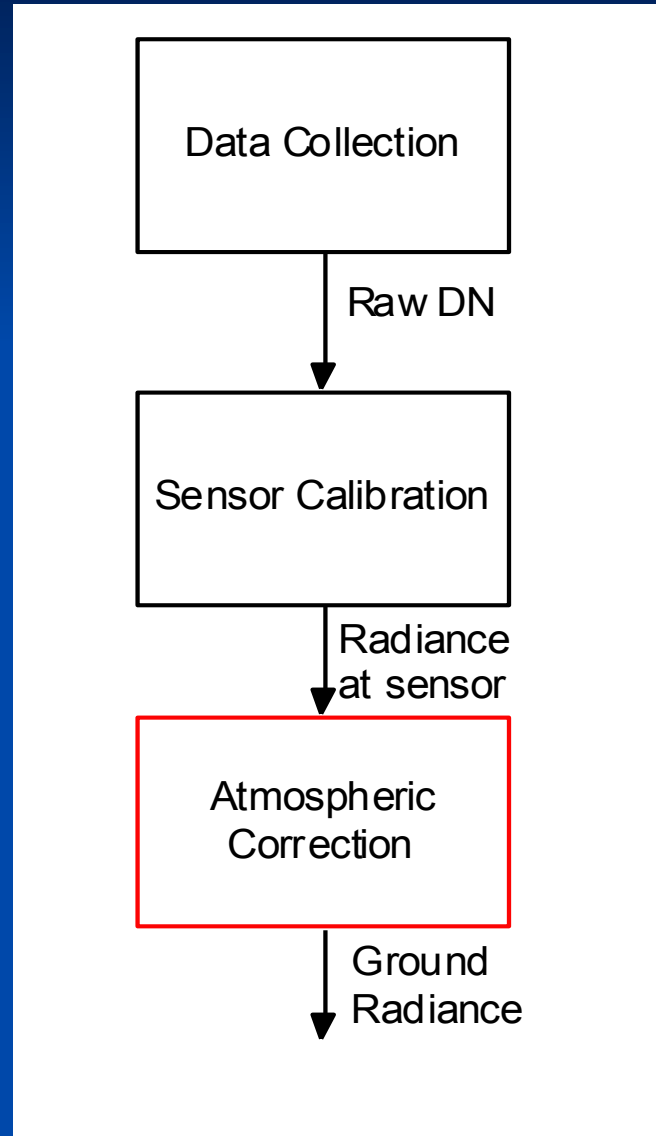
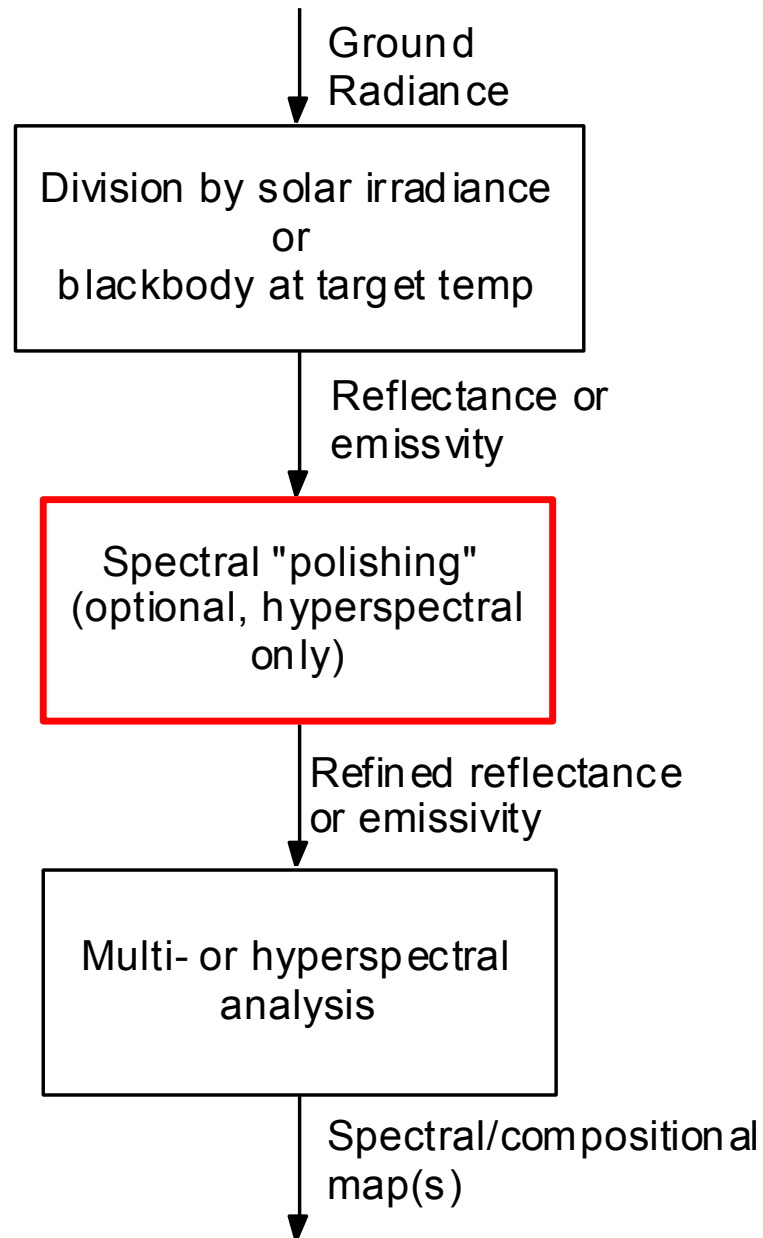
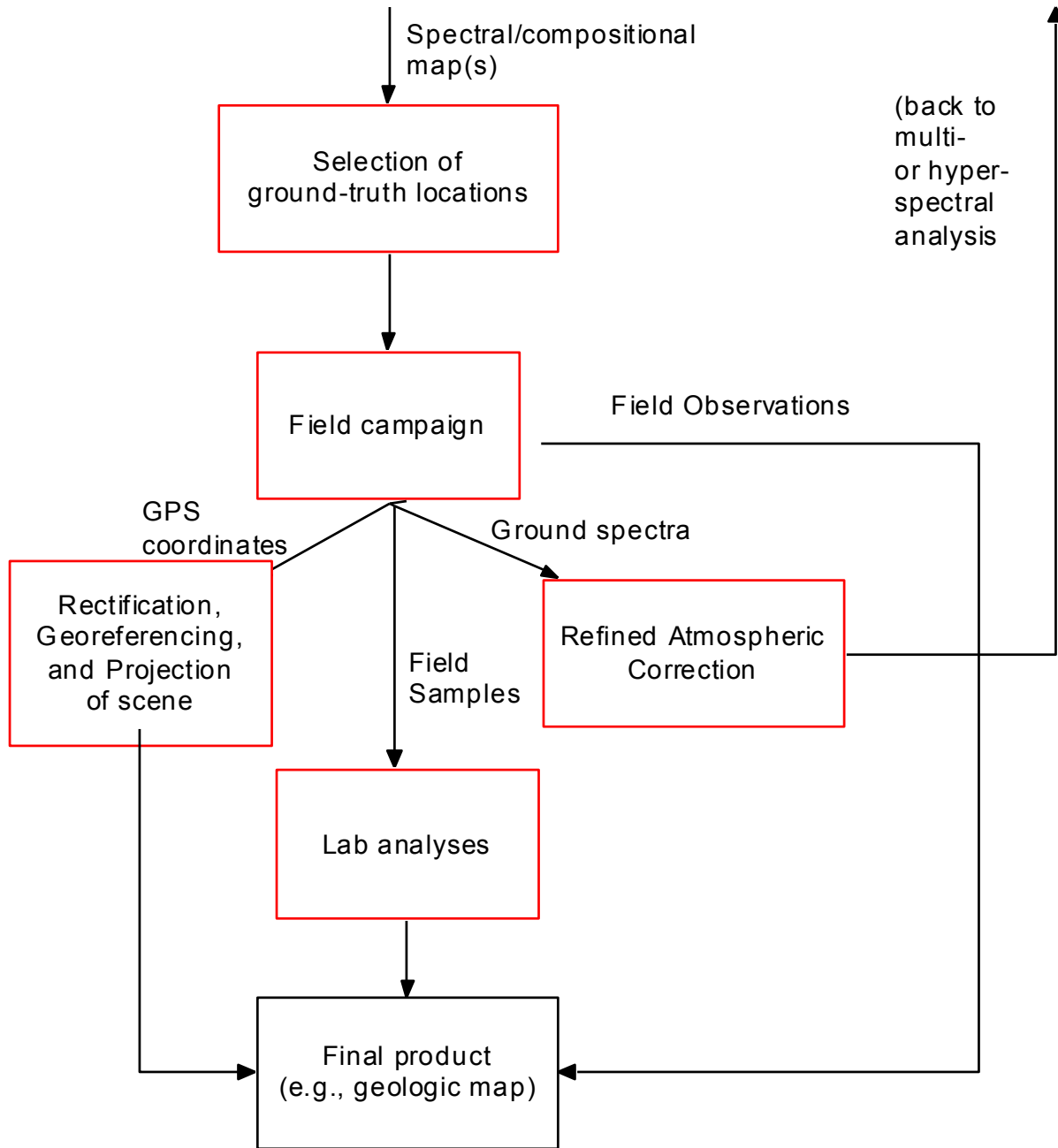


Putting It All Together: A Typical Flowchart for Remote Sensing Projects



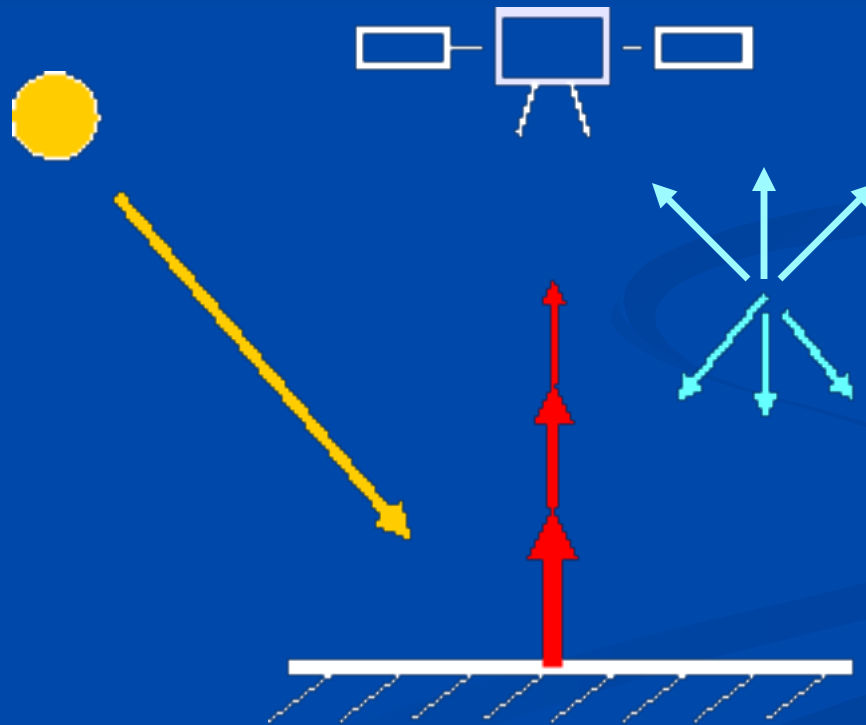




Atmospheric Effects

For measurements of reflected sunlight (Vis/NIR):

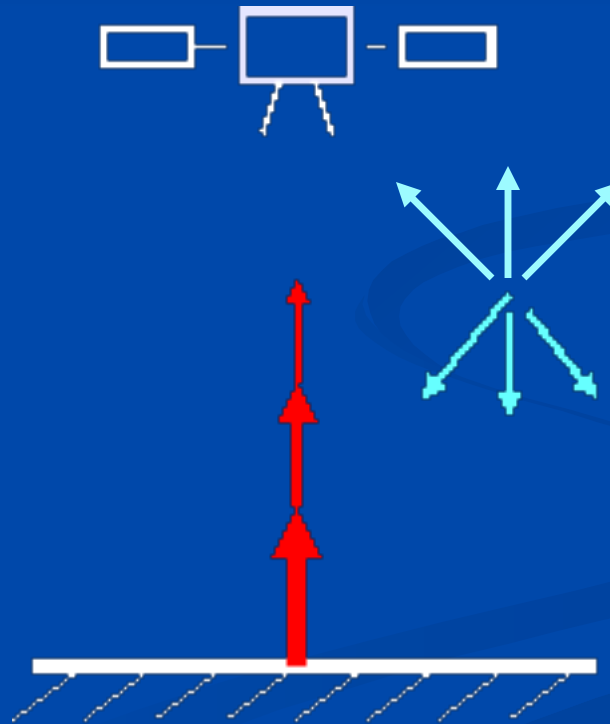
$$I_{sensor} = [RJ_{sun} + RL_{atm_down}] (1 - 2\tau_{atm}) + L_{atm_up}$$



Atmospheric Effects

For measurements of self-emitted light (TIR):

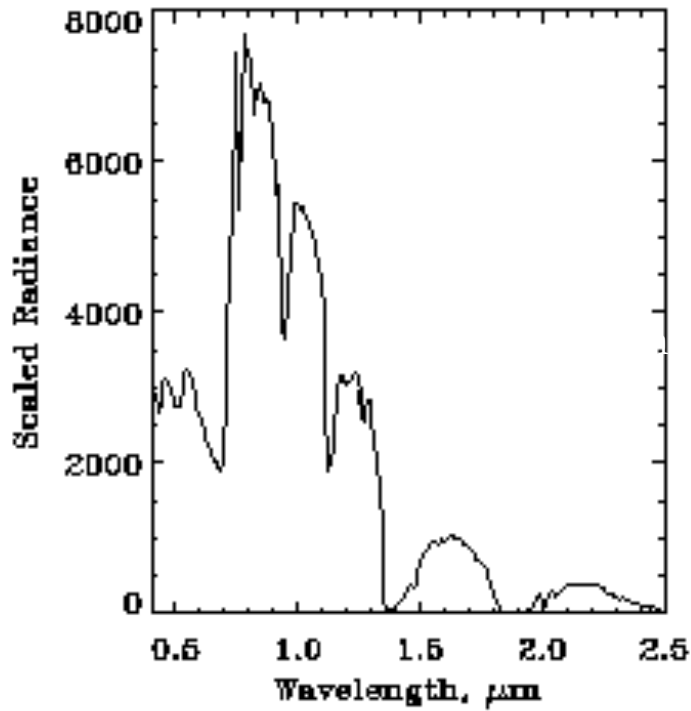
$$I_{sensor} = [\epsilon B_{surface} + RL_{atm_down}] (1 - \tau_{atm}) + L_{atm_up}$$



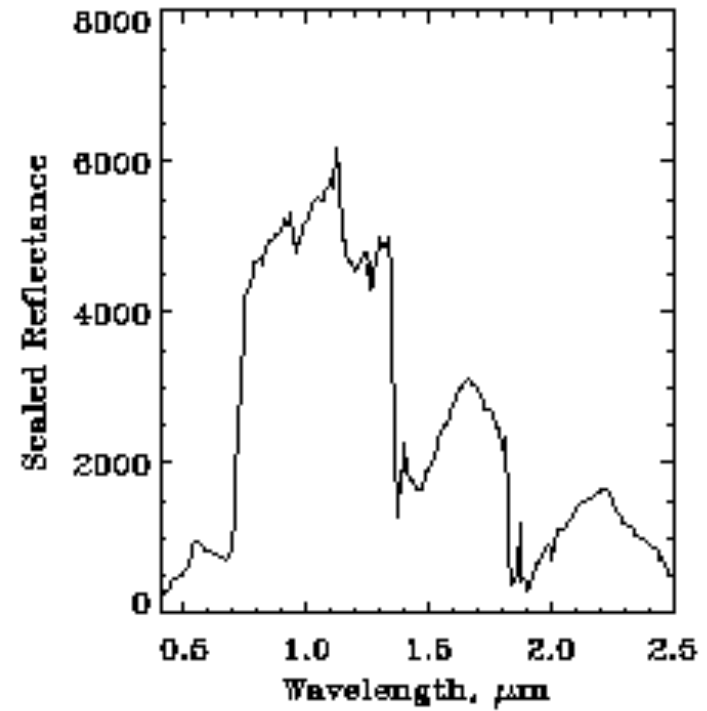
Atmospheric Correction

Model-based corrections:

- Goal is to remove effects of absorption, emission, and scattering of photons by the atmosphere
- For the geologic remote sensing analyst, typically involves use of one of several “black boxes”, e.g. MODTRAN or ATREM.

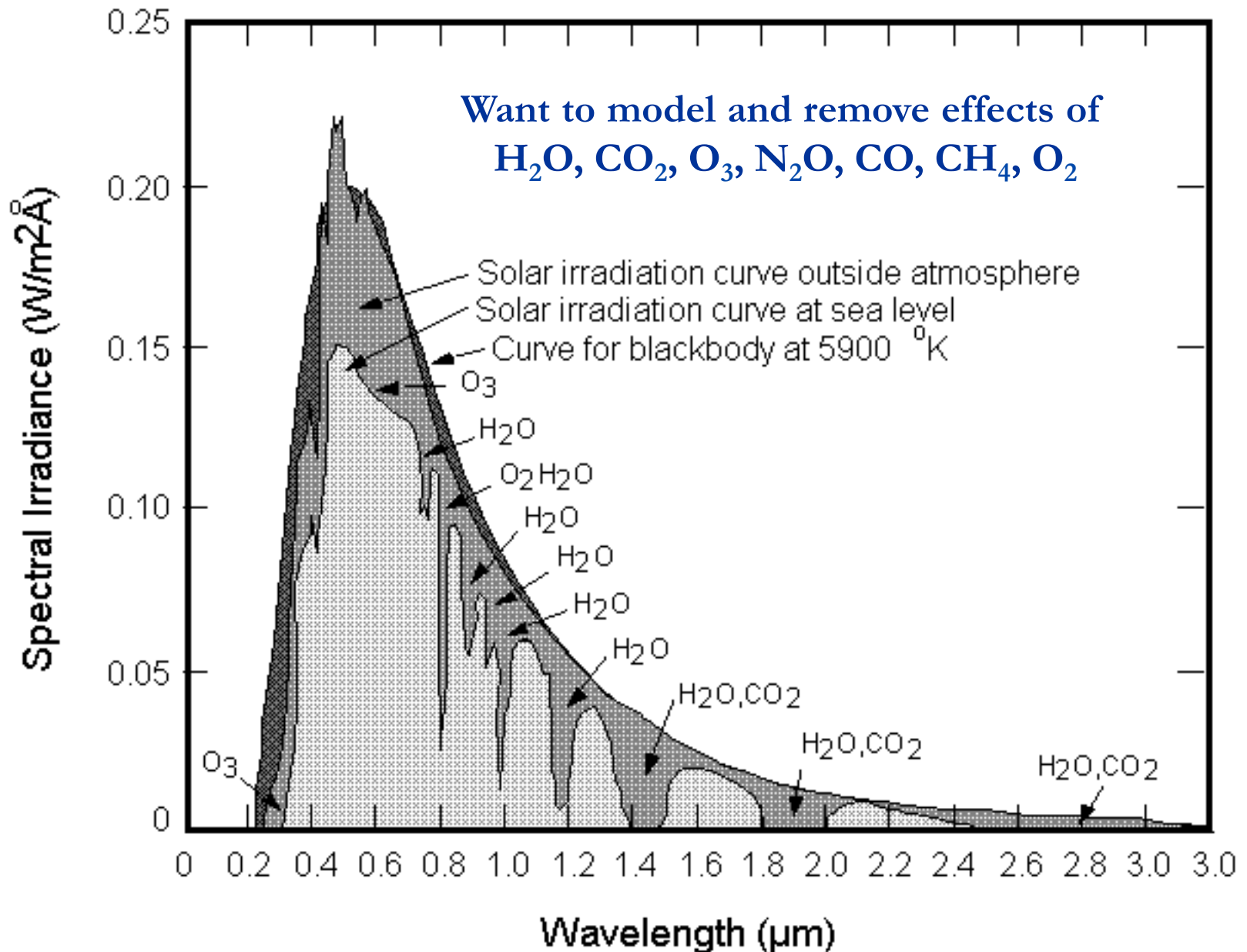


ATREM input



ATREM output

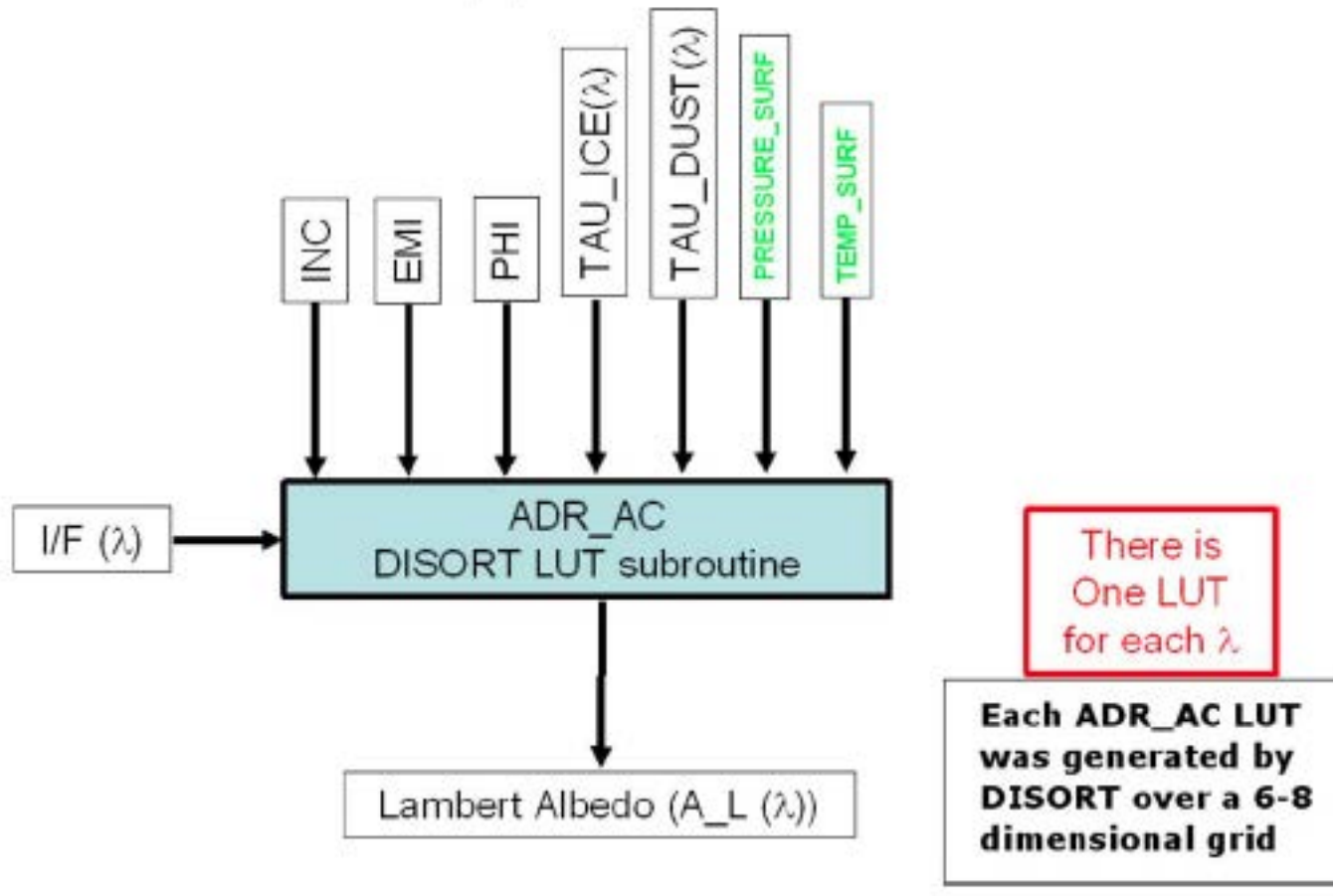
Want to model and remove effects of
 H_2O , CO_2 , O_3 , N_2O , CO , CH_4 , O_2



Solar Spectrum with Atmospheric Absorptions

DIScrete Ordinates Radiative Transfer (DISORT)

Detail of inputs/outputs to the
ADR_AC subroutine



Other types of atmospheric correction:

- IARR (Internal Average Relative Reflectance): Calculates the reflectance of every pixel in the scene relative to the average of all pixels in the scene averaged together.
 - Works best when there are a wide variety of mineralogies in scene, but not great with vegetation.
 - Useful when nothing is known about the scene – e.g., no ground truth spectra, and no model-based atmospheric correction available.
 - Will mute the spectral contrast of components present in a large fraction of the scene.

Other types of atmospheric correction:

- Empirical Line Calibration
 - Employs spectra collected on the ground from known locations in the scene.
 - By comparing pre-correction remote spectrum to ground spectrum from same location, correction values are derived for each wavelength.
 - Works best when multiple locations used, especially if some have low overall reflectance and others have high overall reflectance.
 - Similar in some ways to *volcano scan* technique used on Mars (divide by “atmospheric spectrum” derived from comparing summit and flank of dusty Olympus Mons)

Spectral Polishing

- Goal is to remove any residual atmospheric effects in the spectra
- EFFORT: Empirical Flat Field Optimized Reflectance Transformation:
 - A purely mathematical technique – no physics or geology used.
 - Takes advantage of fact that residual atmospheric effects are usually narrow spectral features, whereas mineralogic features are usually somewhat wider.
 - Fits n-degree polynomial model to spectra from all pixels. For each channel, calculate linear regression of correction factor between data and model. Average correction factors from all pixels

